Nonlinear modulated electrostatic wavepackets in e-p-i plasmas or pair-ion plasmas doped with a stationary charged component

<u>I. Kourakis</u>^{1,2,†}, A. Esfandyari-Kalejahi³, P. K. Shukla²

¹ Universiteit Gent, Sterrenkundig Observatorium

Krijgslaan 281, B-9000 Gent, Belgium

² Institut für Theoretische Physik IV, Fakultät für Physik und Astronomie

Ruhr-Universität Bochum, D-44780 Bochum, Germany

³ Azarbaijan University of Tarbiat Moallem, Faculty of Science

Department of physics, 51745-406, Tabriz, Iran

[†] Email: ioannis@tp4.rub.de; http://www.tp4.rub.de/~ioannis

Abstract

The nonlinear amplitude modulation of electrostatic wave packets propagating in a three-component plasma is investigated, by employing a two-fluid plasma description. Focus is made on electron-positron-ion (e-p-i) plasmas; alternatively, the model describes pair-ion (eg. fullerene) plasmas contaminated by a uniform and stationary minority charged particle species (e.g. defects, or dust particulates). Wave propagation parallel to the external magnetic field is considered.

1. Introduction. Pair plasmas (p.p.), i.e. plasmas consisting of equal mass and opposite charge sign particles, feature properties which do not exist in ordinary (e-i) plasmas. For instance, since the positively and negatively charged particles in p.p. respond on the same frequency scale (unlike electrons and heavy ions), ion-acoustic waves have no counterpart in electron-positron (e-p) plasmas, where the electrostatic (ES) wave dispersion may be of high-frequency parabolic (Langmuir-like) type [1-3], and neither does Faraday rotation. Recently, the production of pair fullerene-ion plasmas in laboratory [3] has enabled experimental studies of pair plasmas rid of intrinsic problems involved in electron-positron plasmas, namely pair recombination processes and strong Landau damping.

In real, e.g. astrophysical contexts, e-p plasmas may be enriched by the additional presence of positive ions. Electron-positron-ion (e-p-i) plasmas appear in the early universe, in active galactic nuclei (AGN) and in pulsar magnetospheres, and may also be created in laboratory (see Refs. in: [4]). Weakly nonlinear low-frequency ES modes in e-p-i plasmas were considered in [5]. Here, we investigate high-frequency oscillations of (light) electrons and positrons (or pair ions) against a neutralizing background of (heavier) ions which, given the frequency range of interest, may be considered immobile.

2. The model. We consider a collisionless plasma, consisting of two inertial species, say 1 and 2, of opposite charge $q_{1/2} = s_{1/2}Zq$ (here $s_1 = -s_2 = +1$) and equal mass $m_{1/2} = m$, and a fixed background of heavier ions (mass m_i , charge $q_i = +Z_ie$); e is the (absolute) electron charge. In specific, 1 and 2 may represent electrons and positrons, in e-p-i plasmas, or heavier C_{60}^+ ions in a pair fullerene-ion plasma, where a minority ion species (e.g. defects) is present.

The (two) inertial dust fluids are described by the moment evolution equations

$$\frac{\partial n_{\alpha}}{\partial t} + \frac{\partial (n_{\alpha} u_{\alpha})}{\partial x} = 0 , \qquad \frac{\partial u_{\alpha}}{\partial t} + u_{\alpha} \frac{\partial u_{\alpha}}{\partial x} = -s_{\alpha} \frac{q}{m} \frac{\partial \phi}{\partial x} - \frac{1}{m n_{\alpha}} \frac{\partial p_{\alpha}}{\partial x} , \tag{1}$$

where t and x are time and (1D) space variables and n_{α} , v_{α} and p_{α} denote the density, velocity and pressure, respectively, of species $\alpha = 1, 2 \equiv +, -$ (of charge sign $s_{\alpha} = \pm 1$). The equation of state $p_{\alpha} = \gamma n_{\alpha} k_B T_{\alpha}$ is assumed to hold, along with $p_{\alpha} = C n_{\alpha}^{\gamma}$; the specific heat ratio $\gamma = (f+2)/f$ (for f degrees of freedom) is here equal to 3; here, T_{α} is the temperature of species α ; k_B is Boltzmann's constant.

The electric potential ϕ obeys Poisson's equation

$$\frac{\partial^2 \phi}{\partial z^2} = -4\pi \sum_{s(pecies)=1}^3 n_s q_s = 4\pi e [Z(n_2 - n_1) - Z_i n_i], \qquad (2)$$

where $n_{1(+)}$ and $n_{2(-)}$ denote the positron (or positive ion) and electron (or negative ion) density, respectively; the background ion density $n_i = n_{i,0}$ is constant. The *rhs* in Eq. (2) cancels at equilibrium, due to the quasi-neutrality condition $Z(n_{2,0} - n_{1,0}) - Z_i n_{i,0} = 0$.

3. Perturbative analysis. The system of (5) Eqs. (1, 2) for the state vector $\mathbf{S} = \{n_1, u_1; n_2, u_2; \phi\}$ supports harmonic electrostatic waves in the form $\mathbf{S} = \hat{\mathbf{S}} \exp[i(k\mathbf{r} - \omega t)] + \text{c.c.}$ In order to study the variation (modulation) of the amplitude(s) \hat{S}_j (here j = 1, ..., 5) due to nonlinearity, we consider small deviations from the equilibrium state $\mathbf{S}^{(0)} = (n_{1,0}, 0; n_{1,0}, 0; 0)^T$, viz. $\mathbf{S} = \mathbf{S}^{(0)} + \epsilon \mathbf{S}^{(1)} + \epsilon^2 \mathbf{S}^{(2)} + ...$, where $\epsilon \ll 1$ is a smallness parameter. We assume $Sj^{(n)} = \sum_{l=-n}^{n} S_{j,l}^{(n)}(X, T) e^{il(k\mathbf{r} - \omega t)}$, for all 5 state variables $(S_{j,-l}^{(n)} = S_{j,l}^{(n)^*}$, for reality), allowing the amplitude(s) to depend on the stretched (slow) coordinates $X = \epsilon(x - v_g t)$ and $T = \epsilon^2 t$; here $v_g = \omega'(k)$ is the wave's group velocity.

The calculation, particularly lengthy yet straightforward, can be found in [4]; also see [6] for details on the method, which essentially implements the generic reductive perturbation method [7] for ES plasma waves.

The (dominant) first harmonic amplitudes are determined (to order $\sim \epsilon^1$) as

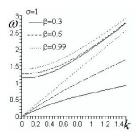
$$n_{+,1}^{(1)} = \frac{n_{-,0}\beta c_s^2 k^2}{\omega^2 - 3\sigma c_s^2 \beta^2 k^2} \frac{Ze\phi_1^{(1)}}{k_B T_-} = \frac{\beta k}{\omega} u_{+,1}^{(1)} \,, \qquad n_{-,1}^{(1)} = -\frac{n_{-,0}c_s^2 k^2}{\omega^2 - 3c_s^2 k^2} \frac{Ze\phi_1^{(1)}}{k_B T_-} = \frac{k}{\omega} u_{-,1}^{(1)} \,,$$

e.g. in terms of the potential correction $\phi_1^{(1)} \equiv \psi$. We have defined the density ratio $\beta = n_{+,0}/n_{-,0}$, the sound speed $c_s = (k_B T_-/m)^{1/2}$, the temperature ratio $\sigma = T_+/T_-$, and the defect- (background ion) density ratio $\delta = n_3/n_2 \equiv n_{i,0}/n_{-,0}$; see that quasineutrality imposes $\beta = 1 - \delta Z_i/Z$, implying $n_+ < n_-$ for positive background ions, i.e. $Z_i > 0$, as implied here (the inverse would hold for $Z_i < 0$); $\beta = 1$ ($\delta = 0$) in p.p. [2].

The linear dispersion relation obtained in $\sim \epsilon^1$ takes the form of a bi-quadratic polynomial equation for ω . Two distinct real solutions are obtained for ω , which for small wave number k values behave as

$$\omega_1 \approx \pm c_{0,L} k \;, \qquad \omega_2 \approx \pm (\omega_g^2 + c_{0,U}^2 k^2)^{1/2} \,,$$
 (3)

where we defined the characteristic speeds $c_{0,L} = c_s[3\beta(1+\sigma\beta)/(1+\beta)]^{1/2}$ and $c_{0,U} = c_s[3(1+\sigma\beta^3)/(1+\beta)]^{1/2}$ and the gap (cutoff) frequency $\omega_g = \omega_{p,-}(1+\beta)^{1/2}$. The L(ower) curve ω_1 is an acoustic branch, while the U(pper) curve ω_2 determines a Langmuir-like optic mode. These results generalize the known dispersion relation for ES modes in pair plasma [1, 3] (here recovered for $\beta = \sigma = 1$). A numerical investigation shows that increasing the fixed ion species density (i.e. decreasing β) results in lower frequency in



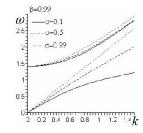


Figure 1: Dispersion relation ω vs. k: effect of variation of β (left) and σ (right).

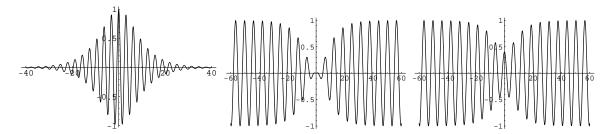


Figure 2: (left) Bright-, (middle) black-, and (right) grey-type soliton solution of (4).

both modes (and, in fact, lower values of $c_{0,U/L}$ and ω_g). On the other hand, for fixed β , decreasing σ results in lower $c_{0,U/L}$ (but does not affect ω_g); see Fig. 1.

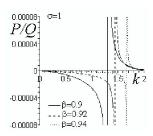
The amplitudes of the 2nd and 0th (constant) harmonic corrections, $S_{j,2}^{(2)}$ and $S_{j,0}^{(2)}$, are obtained in order $\sim \epsilon^2$; the lengthy expressions are omitted for brevity.

4. Nonlinear amplitude evolution equation. In order ϵ^3 , a compatibility condition is obtained, in the form of a nonlinear Schrödinger-type equation (NLSE)

$$i\frac{\partial\psi}{\partial T} + P\frac{\partial^2\psi}{\partial X^2} + Q|\psi|^2\psi = 0, \qquad (4)$$

for the potential correction ψ . Both the dispersion coefficient $P = \omega''(k)/2$ and the nonlinearity coefficient Q, due to carrier wave self-interaction, are lengthy functions of k, σ and β , omitted here, for brevity; the exact expressions can be found in [4]. It may be interesting to trace the behavior of P and Q for long wavelengths, i.e. for $k \ll \lambda_D$ (= $c_s/\omega_{p,-}$). For the lower branch, the coefficients behave as $P \sim -k$ and $Q \sim 1/k$, ensuring modulational stability (since PQ < 0: see in §5 below). For the upper branch, P(k=0) = cst., while $Q \sim k^2$: the optic-type upper mode is therefore unstable.

5. (In)stability profile & envelope localized excitations. The perturbed electric potential is $\psi = \epsilon \psi_0 \cos(kx - \omega t + \Theta) + \mathcal{O}(\epsilon^2)$. It is known [8] that the evolution of a modulated wave whose amplitude obeys Eq. (4) depends on the coefficient product PQ. Eq. (4) supports the plane wave solution $\psi = \psi_0 \exp(iQ|\psi_0|^2T)$; now, perturbing the amplitude as: $\hat{\psi} = \hat{\psi}_0 + \epsilon \hat{\psi}_{1,0} \cos(\tilde{k}X - \tilde{\omega}T)$, one obtains the dispersion relation: $\tilde{\omega}^2 = P \tilde{k}^2 (P\tilde{k}^2 - 2Q|\hat{\psi}_{1,0}|^2)$. If PQ > 0, the amplitude ψ is unstable for $\tilde{k} < \sqrt{2Q/P}|\hat{\psi}_{1,0}|$. If PQ < 0, the amplitude ψ will be stable to external perturbations.



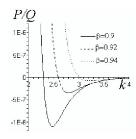


Figure 3: The ratio P/Q vs. k: lower (acoustic) mode β (left); upper mode (right).

This *modulational instability* mechanism is a well known energy localization mechanism in nonlinear dispersive media.

Different localized envelope solutions of Eq. (4) (envelope solitons) exist; see in [6] for a brief outline and analytical expressions (also Refs. therein for details). For PQ > 0, bright envelope modulated wavepackets occur, i.e. localized envelope pulses confining the carrier (see Fig. 2a). For PQ < 0, dark (dark (Fig. 2b) or grey (Fig. 2c) envelope solitons exist, modelling a localized envelope hole (a void) amidst a uniform region.

6. Numerical analysis - conclusions. A numerical analysis shows that both modes are sensitive to variations of the positive-to-negative-ion (or positron-to-electron) density and temperature ratios, β and σ . The lower (acoustic) mode is stable for large wavelengths, and may propagate as a dark-type envelope soliton (a potential dip, or a void). On the other hand, the upper (optic) one is modulationally unstable, and favors the formation of bright-type envelope solitons (pulses). This behavior is depicted in Figs. 3.

These results may be of relevance in experimental [3] and astrophysical [9] contexts. In specific, one may anticipate that the existence of a third minority species in pair plasmas (e.g. defects, or dust) may be used to "tune" the stability of electrostatic modes.

Acknowledgements. I.K. acknowledges partial support by the Deutsche Forschungsgemeinschaft (Bonn, Germany) through the Sonderforschungsbereich (SFB) 591 (*Universelles Verhalten Gleichgewichtsferner Plasmen: Heizung, Transport und Strukturbildung*) Programme.

References

- [1] N. Iwamoto, *Phys. Rev. E* **47**, 604 (1993); G. P. Zank and R. G. Greaves, *Phys. Rev. E* **51**, 6079 (1995).
- [2] I. Kourakis et al., Physics of Plasmas 13 (5), 052117 (2006).
- [3] R. Hatakeyama and W. Oohara, *Phys. Scr.* **116**, 101 (2005); R. Hatakeyama et al., *Proc. ICPP 2006* (Kiev), E010o (2006); also see Refs. therein.
- [4] A. Esfandyari-Kalejahi *et al.*, Electrostatic envelope excitations in e-p-i plasmas, submitted to *J. Phys. A*; online: http://www.tp4.rub.de/~ioannis/publications/2006A11P.pdf.
- [5] M. Salahuddin, H. Saleem and M. Saddiq, Phys. Rev. E 66, 036407 (2002).
- [6] I. Kourakis and P. K. Shukla, Nonlin. Proc. Geophys. 12, 407 (2005).
- [7] T. Taniuti and N. Yajima, *J. Math. Phys.* **10**, 1369 (1969); N. Asano, T. Taniuti and N. Yajima, *J. Math. Phys.* **10**, 2020 (1969).
- [8] Hasegawa, A., Plasma Instabilities and Nonlinear Effects, Springer, 1975; ibid, Optical Solitons in Fibers, Springer, 1989; Remoissenet, M., Waves Called Solitons, Springer, 1994.
- [9] F. C. Michel, Rev. Mod. Phys **54**, 1 (1982).