# RANDOM PARTICLE MOTION IN MAGNETIZED PLASMA 

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#### Abstract

A multivariate Fokker-Planck-type kinetic equation modeling a test - particle weakly interacting with an electrostatic plasma, in the presence of a magnetic field $\mathbf{B}$, is analytically solved in an Ornstein Uhlenbeck - type approximation. A new set of analytic expressions are obtained for variable moments and particle density as a function of time. The process is diffusive.


## 1. Introduction - Formulation of the Problem

The kinetic-theoretical treatment of long-range particle interactions ('collisions') in electrostatic plasma is often based on Landau-type equations [1], describing the evolution of a distribution function (df) $f(\mathbf{v} ; t)$ in velocity space, in the absence of external force fields. This description needs to be modified in the presence of an external field and/or $d f$ spatial inhomogeneity, which not only influence the (free) (Liouville) kinetic operator, but also modify the collision term.

A Fokker-Planck-type kinetic equation (FPE) was recently derived [2,3] from first principles for a test-particle (charge $q$, mass $m$ ) weakly interacting with a plasma embedded in a uniform magnetic field $\mathbf{B}$. This equation, describing the evolution of the $d f$ $f(\mathbf{x}, \mathbf{v} ; t)$ in phase space $\Gamma=\{\mathbf{x}, \mathbf{v}\}$, has the form:

$$
\begin{gather*}
\frac{\partial f}{\partial t}+\mathbf{v} \frac{\partial f}{\partial \mathbf{x}}+\Omega(\mathbf{v} \times \hat{b}) \frac{\partial f}{\partial \mathbf{v}}=\left(\frac{\partial^{2}}{\partial v_{x}^{2}}+\frac{\partial^{2}}{\partial v_{y}^{2}}\right)\left(D_{\perp} f\right)+\frac{\partial^{2}}{\partial v_{z}^{2}}\left(D_{\|} f\right) \\
+2 \Omega^{-1}\left(\frac{\partial^{2}}{\partial v_{x} \partial y}-\frac{\partial^{2}}{\partial v_{y} \partial x}\right)\left(D_{\perp} f\right)+\Omega^{-2}\left(Q+D_{\perp}\right)\left(\frac{\partial^{2}}{\partial x^{2}}+\frac{\partial^{2}}{\partial y^{2}}\right) f \\
-\frac{\partial}{\partial v_{x}}\left(\mathcal{F}_{x} f\right)-\frac{\partial}{\partial v_{y}}\left(\mathcal{F}_{y} f\right)-\frac{\partial}{\partial v_{z}}\left(\mathcal{F}_{z} f\right)+\Omega^{-1} \mathcal{F}_{y} \frac{\partial}{\partial x} f-\Omega^{-1} \mathcal{F}_{x} \frac{\partial}{\partial y} f \tag{1}
\end{gather*}
$$

where $\Omega=q B / m c$ is the cyclotron frequency; $\hat{b}=\mathbf{B} / B$ is the unit vector in the direction of the field $\mathbf{B} ; \partial f / \partial z=0$ by assumption. Note the spatial diffusion term in the righthand side ( $r h s$ ), in fact absent in most previous studies. The lengthy expressions for the coefficients, omitted here, can be found in [3, 4].

In principle, one aims in obtaining an exact solution for $f(\mathbf{x}, \mathbf{v} ; t)$ in order to trace the evolution of variable moments in time, as well as their dependence on physical parameters - the magnetic field $\mathbf{B}$, in particular. However, an exact analytical treatment is not possible, since all coefficients entering the collision term (rhs) are complicated functions of particle velocity $\mathbf{v}$; in addition, they explicitly depend on the magnitude of the external magnetic field. Nevertheless, a numerical study of the coefficients in terms of physical parameters
shows that there exists a region where the diffusion coefficients $D_{\dagger}$ are practically constant (i.e. independent of $\mathbf{v}$ ) while friction terms $\mathcal{F}_{\dagger}$ are linear in $\mathbf{v}$ [3], [4b] (throughout this text, $\dagger$ will denote either $\perp$ or $\|$, referring to quantities perpendicular or parallel, respectively, to $\mathbf{B}$ ). In specific, this is true for low particle velocity value (as compared to the thermal velocity): intuitively speaking, this is close to the standard Langevin picture of a (slow) heavy particle randomly interacting with (faster) light particles surrounding it. This study is devoted to the analytical solution of (1), in the region of validity of this approximation [5].

Setting $D_{\dagger}=$ const., $\mathcal{F}_{\dagger}=\gamma_{\dagger} v_{\dagger}$, eq. (1) may be cast into the standard form of a multivariate (6d) FPE:

$$
\begin{equation*}
\frac{\partial f}{\partial t}=-\sum_{i, j} A_{i j} \frac{\partial}{\partial y_{i}}\left(y_{i} f\right)+\sum_{i, j} D_{i j} \frac{\partial^{2} f}{\partial y_{i} \partial y_{j}} \tag{2}
\end{equation*}
$$

where $f=f(\mathbf{y} ; t) ; \mathbf{y}$ is the position vector $(\mathbf{x}, \mathbf{v})$ in phase space $\Gamma$; cf. (VIII.6.1) in [6]. The diffusion (D) and drift (A) square matrices appearing in (2) are directly derived from (1) via the above assumption and will be omitted here for brevity. Note that $\mathbf{D}$ is symmetric and positive definite. Retain the equilibrium condition: $\gamma_{\dagger}=\frac{m}{T_{\dagger}^{e q}} D_{\dagger}$ $\equiv 2 \beta_{\dagger}^{0} D_{\dagger}$, which is necessary and sufficient in order for the Maxwellian state: $f_{e q}(\mathbf{v})=$ $f_{e q}(\mathbf{0}) e^{-\beta_{\perp}^{0} v_{\perp}^{2}} e^{-\beta_{\|}^{0} v_{\|}^{2}}$ to cancel the rhs in (2). Eq. (1) is now approximated by (2), which defines a multi-dimensional Ornstein-Uhlenbeck process; it may be solved for $f(t)$ via a Green function method. Furthermore, since it describes a Gaussian process, an exact theory exists for the calculation of variable mean values and covariances (see e.g. §VIII. 6 in [6]). The calculation, involving multiple integrations in all $\Gamma$ - space variables ( $\{\mathbf{x}, \mathbf{v}\}$ ), is rather lengthy yet straightforward. This is a brief report of results (exposed in [3] in detail).

## 2. Exact Solution for a Maxwelian Initial Velocity Distribution

Assuming a Maxwellian initial distribution of the form:

$$
\begin{equation*}
f_{0}(\mathbf{x}, \mathbf{v})=\frac{1}{\pi^{3 / 2}} \beta_{\perp} \beta_{\|}^{1 / 2} e^{-\beta_{\perp} v_{\perp}^{2}} e^{-\beta_{\|} v_{\|}^{2}} \delta(\mathbf{x}) \tag{3}
\end{equation*}
$$

( $\beta_{\dagger}=m / 2 T_{\dagger}$ ) we obtain the (time-dependent) $d f$ :

$$
\begin{equation*}
f(\mathbf{x}, \mathbf{v} ; t)=\frac{\tilde{\beta}_{\perp} \tilde{\beta}_{\|}^{1 / 2} \tilde{\beta}_{\perp}^{(X)}}{\pi^{5 / 2}} e^{-\tilde{\beta}_{\perp} v_{\perp}^{2}} e^{-\tilde{\beta}_{\|} v_{\|}^{2}} e^{-\tilde{\beta}_{\perp}^{(X)}} \Xi^{2} I_{0}\left(4 \frac{\tilde{\beta}_{\perp}^{(X)} \tilde{\beta}_{\perp} \xi}{\Omega} v_{\perp} \rho\right) \tag{4}
\end{equation*}
$$

where $\rho=x^{2}+y^{2}, v_{\perp}=v_{x}^{2}+v_{y}^{2}$ and $v_{\|}=v_{z}$; note the definitions:

$$
\begin{align*}
\tilde{\beta}_{\dagger}(t)= & \frac{\theta}{\left(1-e^{-2 \gamma_{+} t}\right) \theta+e^{-2 \gamma_{\dagger} t}} \beta_{\dagger}^{0}, \quad \Xi(\tau)=\rho^{2}+\left(\frac{2 \tilde{\beta}_{\perp}(\tau) \xi(\tau)}{\Omega} v_{\perp}\right)^{2} \\
\tilde{\beta}_{\perp}^{(X)}(\tau)= & \Omega^{2} \beta_{\perp}^{0}\left\{1-e^{-2 \tau}+\frac{\left(1-e^{-\tau}\right)^{2}}{\theta}\right. \\
& \left.\quad-\frac{1}{\theta} \frac{\theta^{2}\left(1-e^{-2 \tau}\right)^{2}+e^{-2 \tau}\left(1-e^{-\tau}\right)^{2}}{\theta\left(1-e^{-2 \tau}\right)+e^{-2 \tau}}+4 \beta_{\perp}^{0} Q t\right\}^{-1} \\
\xi(\tau)= & \frac{1}{\beta_{\perp}^{0}} \frac{1}{2 \theta}\left\{\theta^{2}\left(1-e^{-2 \tau}\right)^{2}+e^{-2 \tau}\left(1-e^{-\tau}\right)^{2}\right\}^{1 / 2} \tag{5}
\end{align*}
$$

where $\theta=\beta_{\dagger} / \beta_{\dagger}^{0}=T_{\dagger}^{e q} / T_{\dagger}, \tau=\gamma_{\perp} t$. We see that the velocity distribution will relax to the equilibrium state anticipated above, as physically expected, since: $\lim _{\tau \rightarrow \infty} \tilde{\beta}_{\dagger}(\tau)$ $=\frac{\gamma_{\dagger}}{2 D_{\dagger}} \equiv \beta_{\dagger}^{0}=\frac{m}{2 T_{\dagger}^{e q}}$, while spatial distribution will exhibit a classical diffusive behaviour, under the influence of collisions; check that, for $\tau \gg \gamma_{\dagger}^{-1}: \tilde{\beta}_{\perp}^{(X)}(\tau)$ behaves as $\approx \Omega^{2} \beta_{\perp}^{0} /\left(1+4 \beta_{\perp}^{0} Q t\right)$. A similar qualitative behaviour is also obtained for a Maxwellian (i.e. not localized, cf. (3)) spatial distribution at $t=0$, and also for an initial velocity distribution of the type: $f_{\perp}(\mathbf{v} ; t=0)=\delta\left(\mathbf{v}_{\perp}-\mathbf{v}_{\|}^{0}\right) \delta\left(v_{\|}-v_{\|}^{0}\right)$ (cf. fig. 1b); details are omitted here, for lack of space [3].

( a )

(b)

Fig. 1. (a) The inverse temperature $\tilde{\beta}_{\dagger}(t)$ (scaled over $\tilde{\beta}_{\dagger}(\infty)=\beta_{\dagger}^{0}$ ) versus time $\tau=\gamma_{\dagger} t$ for different values of $\theta=\beta / \beta^{0}=T_{e q} / T(0)$. Collisions appear more efficient in accelerating initially slow particles $(\theta>1)$ than slowing down fast ones $(\theta<1)$.
(b) The evolution $f_{\perp}\left(v_{\|} ; t\right)$ of an initial $\delta\left(v_{\|}-v_{\|}^{0}\right)$ state (sharp solid line) (\| -part; see in §2) versus velocity $v_{\|}\left(v_{\|} \in \Re\right)$ (normalized over the thermal velocity $\beta_{\|}^{0-1 / 2}$ ) for an initial velocity value of $2 v_{t h}$. The initial sharp profile spreads fast and attains the final zero-averaged thermalized state within a few time constants.

## 3. Density Profile - Moments

The mean particle density: $n(\mathbf{x}, t)=\int d \mathbf{v} f(\mathbf{x}, \mathbf{v} ; t)$ may now be calculated. Considering the initial condition: $f_{0}(\mathbf{x}, \mathbf{v})=\delta\left(\mathbf{x}-\mathbf{x}_{0}\right) \delta\left(\mathbf{v}-\mathbf{v}_{0}\right)$, we obtain:

$$
\begin{equation*}
n(\rho, t)=\frac{1}{2 \pi} \frac{1}{L^{2}(t)} e^{-\rho^{2} / 2 L^{2}(t)} e^{-R_{0}^{2}(t) / 2 L^{2}(t)} I_{0}\left(\frac{R_{0}(t) \rho}{L^{2}(t)}\right) \tag{6}
\end{equation*}
$$

(for simplicity the direction $\|$ to the field was neglected and the gyro-phase $\phi$ was averaged out); $I_{0}(x)$ denotes the modified Bessel function; the mean square displacement $L^{2}$ reads:

$$
\begin{equation*}
L^{2}(t)=\Omega^{-2}\left[\frac{T_{\perp}^{e q}}{m}\left(1-e^{-2 \gamma_{\perp} t}\right)+4 Q t\right] \tag{7}
\end{equation*}
$$

and $R_{0}(t)=\left(1+e^{-2 \gamma_{\perp} t}-2 e^{-\gamma_{\perp} t} \cos \Omega t\right)^{1 / 2} v_{\perp}^{0} / \Omega$; see that $R_{0} \approx \frac{v_{\perp}^{0}}{\Omega} \equiv \rho_{L}^{0}$ after a while. For zero initial velocity, $R_{0}=0$ so a purely Gaussian profile is recovered. For higher $v_{\perp}^{0}$, the distribution spreads in space, as expected. See that $L^{2}(t)$ asymptotically behaves as $\sim t$, hence the classical diffusion mechanism encountered in $\S 2$. Earlier results predicting diffusion as $\sim B^{-2}$ are thus confirmed [7].

Finally, the (symmetric) covariance matrix $\left\langle\left\langle y_{k} y_{l}\right\rangle\right\rangle=\left\langle y_{k} y_{l}\right\rangle-\left\langle y_{k}\right\rangle\left\langle y_{l}\right\rangle$ (where $\left.\mathbf{y}=\left(x, y, z ; v_{x}, v_{y}, v_{z}\right)\right)$ can be evaluated via the same formalism. The results for all


Fig. 2. The evolution $f_{\perp}\left(v_{\perp} ; t\right)$ of an initial Maxwellian state ( $\perp$-part, as given by (5), (6)) versus velocity $v_{\perp}$ (normalized over the thermal velocity $\beta_{\perp}^{0-1 / 2}$ ) for two values of $\theta=\beta_{*} / \beta_{*}^{0}=$ $T_{e q} / T_{*}(0):$ (a) $\theta=0.01$ (high initial plasma temperature, see black solid line at the bottom) and (b) $\theta=10$ (low initial plasma temperature, see black solid line on top). In the second case, the initial distribution relaxes faster to the final equilibrium state (practically attained after 2 time constants) in agreement with Fig. 1a.
elements are obtained in [3] and briefly exposed in [5]. For velocity covariances we obtain: $\left\langle\left\langle v_{x} v_{x}\right\rangle\right\rangle=\left\langle\left\langle v_{y} v_{y}\right\rangle\right\rangle=\frac{T_{\perp}^{e q}}{m}\left(1-e^{-2 \gamma_{\perp} t}\right)$ (also for $\left\langle\left\langle v_{z} v_{z}\right\rangle\right\rangle$, upon $\perp \rightarrow \|$ ). In the plane $\perp \mathbf{B}:\langle\langle x x\rangle\rangle=\langle\langle y y\rangle\rangle)=L^{2}$; cf. (7) above.

In conclusion, the analytical treatment of random electrostatic interactions modelled by the Fokker-Planck-type equation (1) in an Ornstein-type approximation reveals a classical diffusive behaviour in space and allows for an exact study of plasma relaxation as well as moment evolution in time.

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## References

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[1] R. Balescu, Transport Processes in Plasmas, vol. 1, Classical Transport North Holland, Amsterdam (1988).
[2] I. Kourakis, Plasma Phys. Control. Fusion 41587 (1999); also, A. Grecos \& I. Kourakis, in preparation.
[3] I. Kourakis, PhD thesis, U.L.B., Brussels, Belgium (2002).
[4] (a) I.Kourakis, D.Carati, B.Weyssow, Proc. Int. Conf. Plasma Phys., Québec (2000), 49-53;
(b) I.Kourakis, European Conference Abstracts (ECA) Vol. 26B, P-4.008 (2002).
[5] I. Kourakis \& A. Grecos, Comm. Nonlin. Sci. Num. Sim., 8 (3-4) (2003), in press.
[6] N. G. Van-Kampen, Stochastic Processes, North-Holland, Amsterdam (1992).
[7] M. Rosenbluth \& A. Kaufman, Phys. Rev. 109 (1) (1958), 1; J. Taylor, Phys. Rev. Lett. 6 (6) (1961), 262.

