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# **RANDOM PARTICLE MOTION IN MAGNETIZED PLASMA**

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A multivariate Fokker-Planck-type kinetic equation modeling a test - particle weakly interacting with an electrostatic plasma, in the presence of a magnetic field **B**, is analytically solved in an Ornstein - Uhlenbeck - type approximation. A new set of analytic expressions are obtained for variable moments and particle density as a function of time. The process is diffusive.

# 1. Introduction - Formulation of the Problem

The kinetic-theoretical treatment of long-range particle interactions ('collisions') in electrostatic plasma is often based on Landau-type equations [1], describing the evolution of a distribution function  $(df) f(\mathbf{v}; t)$  in velocity space, in the absence of external force fields. This description needs to be modified in the presence of an external field and/or df spatial inhomogeneity, which not only influence the (free) (Liouville) kinetic operator, but also modify the collision term.

A Fokker-Planck-type kinetic equation (*FPE*) was recently derived [2, 3] from first principles for a test-particle (charge q, mass m) weakly interacting with a plasma embedded in a uniform magnetic field **B**. This equation, describing the evolution of the  $df f(\mathbf{x}, \mathbf{v}; t)$  in phase space  $\Gamma = {\mathbf{x}, \mathbf{v}}$ , has the form:

$$\frac{\partial f}{\partial t} + \mathbf{v} \frac{\partial f}{\partial \mathbf{x}} + \Omega \left( \mathbf{v} \times \hat{b} \right) \frac{\partial f}{\partial \mathbf{v}} = \left( \frac{\partial^2}{\partial v_x^2} + \frac{\partial^2}{\partial v_y^2} \right) \left( D_\perp f \right) + \frac{\partial^2}{\partial v_z^2} \left( D_\parallel f \right) \\
+ 2 \Omega^{-1} \left( \frac{\partial^2}{\partial v_x \partial y} - \frac{\partial^2}{\partial v_y \partial x} \right) \left( D_\perp f \right) + \Omega^{-2} \left( Q + D_\perp \right) \left( \frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} \right) f \\
- \frac{\partial}{\partial v_x} \left( \mathcal{F}_x f \right) - \frac{\partial}{\partial v_y} \left( \mathcal{F}_y f \right) - \frac{\partial}{\partial v_z} \left( \mathcal{F}_z f \right) + \Omega^{-1} \mathcal{F}_y \frac{\partial}{\partial x} f - \Omega^{-1} \mathcal{F}_x \frac{\partial}{\partial y} f \qquad (1)$$

where  $\Omega = qB/mc$  is the cyclotron frequency;  $\hat{b} = \mathbf{B}/B$  is the unit vector in the direction of the field **B**;  $\partial f/\partial z = 0$  by assumption. Note the *spatial diffusion* term in the righthand side (*rhs*), in fact absent in most previous studies. The lengthy expressions for the coefficients, omitted here, can be found in [3, 4].

In principle, one aims in obtaining an exact solution for  $f(\mathbf{x}, \mathbf{v}; t)$  in order to trace the evolution of variable moments in time, as well as their dependence on physical parameters - the magnetic field **B**, in particular. However, an exact analytical treatment is not possible, since all coefficients entering the collision term (*rhs*) are complicated functions of particle velocity  $\mathbf{v}$ ; in addition, they explicitly depend on the magnitude of the external magnetic field. Nevertheless, a numerical study of the coefficients in terms of physical parameters

shows that there exists a region where the diffusion coefficients  $D_{\dagger}$  are practically *constant* (i.e. independent of **v**) while friction terms  $\mathcal{F}_{\dagger}$  are *linear* in **v** [3], [4b] (throughout this text,  $\dagger$  will denote either $\perp$  or  $\parallel$ , referring to quantities perpendicular or parallel, respectively, to **B**). In specific, this is true for low particle velocity value (as compared to the thermal velocity): intuitively speaking, this is close to the standard Langevin picture of a (slow) heavy particle randomly interacting with (faster) light particles surrounding it. This study is devoted to the analytical solution of (1), in the region of validity of this approximation [5].

Setting  $D_{\dagger} = const.$ ,  $\mathcal{F}_{\dagger} = \gamma_{\dagger} v_{\dagger}$ , eq. (1) may be cast into the standard form of a multivariate (6d) *FPE*:

$$\frac{\partial f}{\partial t} = -\sum_{i,j} A_{ij} \frac{\partial}{\partial y_i} (y_i f) + \sum_{i,j} D_{ij} \frac{\partial^2 f}{\partial y_i \partial y_j}$$
(2)

where  $f = f(\mathbf{y}; t)$ ;  $\mathbf{y}$  is the position vector  $(\mathbf{x}, \mathbf{v})$  in phase space  $\Gamma$ ; cf. (VIII.6.1) in [6]. The diffusion (**D**) and drift (**A**) square matrices appearing in (2) are directly derived from (1) via the above assumption and will be omitted here for brevity. Note that **D** is symmetric and positive definite. Retain the equilibrium condition:  $\gamma_{\dagger} = \frac{m}{T_{\dagger}^{eq}}D_{\dagger}$  $\equiv 2\beta_{\dagger}^{0}D_{\dagger}$ , which is necessary and sufficient in order for the Maxwellian state:  $f_{eq}(\mathbf{v}) = f_{eq}(\mathbf{0}) e^{-\beta_{\perp}^{0}v_{\perp}^{2}} e^{-\beta_{\parallel}^{0}v_{\parallel}^{2}}$  to cancel the *rhs* in (2). Eq. (1) is now approximated by (2), which defines a multi-dimensional Ornstein-Uhlenbeck process; it may be solved for f(t) via a Green function method. Furthermore, since it describes a Gaussian process, an exact theory exists for the calculation of variable mean values and covariances (see e.g. §VIII.6 in [6]). The calculation, involving multiple integrations in all  $\Gamma$ - space variables ({**x**, **v**}), is rather lengthy yet straightforward. This is a brief report of results (exposed in [3] in detail).

### 2. Exact Solution for a Maxwelian Initial Velocity Distribution

Assuming a Maxwellian initial distribution of the form:

$$f_0(\mathbf{x}, \mathbf{v}) = \frac{1}{\pi^{3/2}} \beta_{\perp} \beta_{\parallel}^{1/2} e^{-\beta_{\perp} v_{\perp}^2} e^{-\beta_{\parallel} v_{\parallel}^2} \delta(\mathbf{x})$$
(3)

 $(\beta_{\dagger} = m/2T_{\dagger})$  we obtain the (time-dependent) df:

$$f(\mathbf{x}, \mathbf{v}; t) = \frac{\tilde{\beta}_{\perp} \tilde{\beta}_{\parallel}^{1/2} \tilde{\beta}_{\perp}^{(X)}}{\pi^{5/2}} e^{-\tilde{\beta}_{\perp} v_{\perp}^{2}} e^{-\tilde{\beta}_{\parallel} v_{\parallel}^{2}} e^{-\tilde{\beta}_{\perp}^{(X)} \Xi^{2}} I_{0} \left( 4 \frac{\tilde{\beta}_{\perp}^{(X)} \tilde{\beta}_{\perp} \xi}{\Omega} v_{\perp} \rho \right)$$
(4)

where  $\rho = x^2 + y^2$ ,  $v_{\perp} = v_x^2 + v_y^2$  and  $v_{\parallel} = v_z$ ; note the definitions:

$$\begin{split} \tilde{\beta}_{\dagger}(t) &= \frac{\theta}{(1 - e^{-2\gamma_{\dagger}t})\theta + e^{-2\gamma_{\dagger}t}}\beta_{\dagger}^{0}, \qquad \Xi(\tau) = \rho^{2} + \left(\frac{2\tilde{\beta}_{\perp}(\tau)\xi(\tau)}{\Omega}v_{\perp}\right)^{2} \\ \tilde{\beta}_{\perp}^{(X)}(\tau) &= \Omega^{2}\beta_{\perp}^{0}\left\{1 - e^{-2\tau} + \frac{(1 - e^{-\tau})^{2}}{\theta} - \frac{1}{\theta}\frac{\theta^{2}(1 - e^{-2\tau})^{2} + e^{-2\tau}(1 - e^{-\tau})^{2}}{\theta(1 - e^{-2\tau}) + e^{-2\tau}} + 4\beta_{\perp}^{0}Qt\right\}^{-1} \\ \xi(\tau) &= \frac{1}{\beta_{\perp}^{0}}\frac{1}{2\theta}\left\{\theta^{2}(1 - e^{-2\tau})^{2} + e^{-2\tau}(1 - e^{-\tau})^{2}\right\}^{1/2} \end{split}$$
(5)

where  $\theta = \beta_{\dagger}/\beta_{\dagger}^{0} = T_{\dagger}^{eq}/T_{\dagger}$ ,  $\tau = \gamma_{\perp}t$ . We see that the velocity distribution will relax to the equilibrium state anticipated above, as physically expected, since:  $\lim_{\tau \to \infty} \tilde{\beta}_{\dagger}(\tau)$  $= \frac{\gamma_{\dagger}}{2D_{\dagger}} \equiv \beta_{\dagger}^{0} = \frac{m}{2T_{\dagger}^{eq}}$ , while spatial distribution will exhibit a *classical* diffusive behaviour, under the influence of collisions; check that, for  $\tau \gg \gamma_{\dagger}^{-1}$ :  $\tilde{\beta}_{\perp}^{(X)}(\tau)$  behaves as  $\approx \Omega^{2} \beta_{\perp}^{0}/(1+4\beta_{\perp}^{0} Q t)$ . A similar qualitative behaviour is also obtained for a Maxwellian (i.e. not localized, cf. (3)) spatial distribution at t = 0, and also for an initial velocity distribution of the type:  $f_{\perp}(\mathbf{v}; t = 0) = \delta(\mathbf{v}_{\perp} - \mathbf{v}_{\parallel}^{0}) \delta(v_{\parallel} - v_{\parallel}^{0})$  (cf. fig. 1b); details are omitted here, for lack of space [3].



Fig. 1. (a) The inverse temperature  $\tilde{\beta}_{\dagger}(t)$  (scaled over  $\tilde{\beta}_{\dagger}(\infty) = \beta^{0}_{\dagger}$ ) versus time  $\tau = \gamma_{\dagger} t$  for different values of  $\theta = \beta/\beta^{0} = T_{eq}/T(0)$ . Collisions appear more efficient in accelerating initially slow particles ( $\theta > 1$ ) than slowing down fast ones ( $\theta < 1$ ).

(b) The evolution  $f_{\perp}(v_{\parallel}; t)$  of an initial  $\delta(v_{\parallel} - v_{\parallel}^{0})$  state (sharp solid line) ( $\parallel$  -part; see in §2) versus velocity  $v_{\parallel}$  ( $v_{\parallel} \in \Re$ ) (normalized over the thermal velocity  $\beta_{\parallel}^{0^{-1/2}}$ ) for an initial velocity value of  $2v_{th}$ . The initial sharp profile spreads fast and attains the final zero-averaged thermalized state within a few time constants.

#### 3. Density Profile - Moments

The mean particle density:  $n(\mathbf{x}, t) = \int d\mathbf{v} f(\mathbf{x}, \mathbf{v}; t)$  may now be calculated. Considering the initial condition:  $f_0(\mathbf{x}, \mathbf{v}) = \delta(\mathbf{x} - \mathbf{x_0}) \delta(\mathbf{v} - \mathbf{v_0})$ , we obtain:

$$n(\rho, t) = \frac{1}{2\pi} \frac{1}{L^2(t)} e^{-\rho^2/2L^2(t)} e^{-R_0^2(t)/2L^2(t)} I_0\left(\frac{R_0(t)\rho}{L^2(t)}\right)$$
(6)

(for simplicity the direction || to the field was neglected and the gyro-phase  $\phi$  was averaged out);  $I_0(x)$  denotes the modified Bessel function; the mean square displacement  $L^2$  reads:

$$L^{2}(t) = \Omega^{-2} \left[ \frac{T_{\perp}^{eq}}{m} (1 - e^{-2\gamma_{\perp} t}) + 4Qt \right]$$
(7)

and  $R_0(t) = (1 + e^{-2\gamma_{\perp}t} - 2e^{-\gamma_{\perp}t} \cos \Omega t)^{1/2} v_{\perp}^0 / \Omega$ ; see that  $R_0 \approx \frac{v_{\perp}^0}{\Omega} \equiv \rho_L^0$  after a while. For zero initial velocity,  $R_0 = 0$  so a purely Gaussian profile is recovered. For higher  $v_{\perp}^0$ , the distribution spreads in space, as expected. See that  $L^2(t)$  asymptotically behaves as  $\sim t$ , hence the classical diffusion mechanism encountered in §2. Earlier results predicting diffusion as  $\sim B^{-2}$  are thus confirmed [7].

Finally, the (symmetric) covariance matrix  $\langle \langle y_k y_l \rangle \rangle = \langle y_k y_l \rangle - \langle y_k \rangle \langle y_l \rangle$  (where  $\mathbf{y} = (x, y, z; v_x, v_y, v_z)$ ) can be evaluated via the same formalism. The results for all



Fig. 2. The evolution  $f_{\perp}(v_{\perp}; t)$  of an initial Maxwellian state  $(\perp -part, as given by (5), (6))$  versus velocity  $v_{\perp}$  (normalized over the thermal velocity  $\beta_{\perp}^{0^{-1/2}}$ ) for two values of  $\theta = \beta_*/\beta_*^0 = T_{eq}/T_*(0)$ : (a)  $\theta = 0.01$  (high initial plasma temperature, see black solid line at the bottom) and (b)  $\theta = 10$  (low initial plasma temperature, see black solid line on top). In the second case, the initial distribution relaxes faster to the final equilibrium state (practically attained after 2 time constants) in agreement with Fig. 1a.

elements are obtained in [3] and briefly exposed in [5]. For velocity covariances we obtain:  $\langle \langle v_x v_x \rangle \rangle = \langle \langle v_y v_y \rangle \rangle = \frac{T_{\perp}^{eq}}{m} (1 - e^{-2\gamma_{\perp} t})$  (also for  $\langle \langle v_z v_z \rangle \rangle$ , upon  $\perp \rightarrow \parallel$ ). In the plane  $\perp \mathbf{B}: \langle \langle x x \rangle \rangle = \langle \langle y y \rangle \rangle = L^2$ ; cf. (7) above.

In conclusion, the analytical treatment of random electrostatic interactions modelled by the Fokker-Planck-type equation (1) in an Ornstein-type approximation reveals a classical diffusive behaviour in space and allows for an exact study of plasma relaxation as well as moment evolution in time.

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- 174 -